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Today Solitons.

Ex 1. O(1) model with spontaneous symmetry breaking.

$$\mathcal{L} = \frac{1}{2} \partial_\mu \phi \partial^\mu \phi - \frac{1}{2} \lambda (\phi^2 - c^2)^2, \quad c > 0$$

2 vacua $\phi_0 = c, \phi_0 = -c$.

Now chose number of space time dimeensions $d = 1 + 1$.

$$\mathcal{L} = \frac{1}{2} (\dot{\phi}^2 - \phi'^2) - \frac{1}{2} \lambda (\phi^2 - c^2)^2$$

what happens if $x \rightarrow +\infty$ and $x \rightarrow -\infty$ choose different vacua.

$$\phi_s(x) \rightarrow +c, \quad x \rightarrow +\infty$$

$$\phi_s(x) \rightarrow -c, \quad x \rightarrow -\infty$$

Then $\phi_s(x)$ is a static solution to equations minimizing potential energy U .

$$U = \int dx \left(\frac{1}{2} \phi'^2 + \frac{1}{2} \lambda (\phi^2 - c^2)^2 \right) = \int dx \mathcal{L}(\phi, \phi')$$

Equation: $\partial_x^2 \phi_s = \lambda \phi (\phi^2 - c^2)$.

No explicit x dependency in $U \Rightarrow$ Jacobi integral conserved.

$$J = \frac{d}{dx} \frac{\partial L}{\partial \phi'} - \frac{\partial L}{\partial \phi} = \frac{1}{2} \phi'^2 - \frac{1}{2} \lambda (\phi^2 - c^2)^2 = \text{constant} = 0$$

$$\phi' = \pm \sqrt{\lambda} (c^2 - \phi^2)$$

$$\frac{\phi'}{c} = \sqrt{\lambda} c \left(1 - \frac{\phi^2}{c^2} \right)$$

$$\frac{\phi}{c} = \tanh(\sqrt{\lambda} c (x - x_0))$$

This configuration is called soliton, energy localized to region of width $\Delta x = 1/\sqrt{\lambda} c$ around $x = x_0$. Total energy = soliton mass $\sim \sqrt{\lambda} c^3$.

$\phi_s(x - x_0)$ describes soliton (extended particle) at rest at x_0 . This soliton at rest, but the theory is relativistic \Rightarrow by Lorentz transformation you get moving solution with constant velocity v and relativistic energy momentum relation $E^2 = m^2 + p^2, v = p/E$. This type of soliton could occur as a domain wall in astrophysics.

Solitons in $d > 1 + 1$?

O(n) model in $d - 1$ space dimensions.

$n = 2$. O(2) model in $d = 2 + 1$, 2 space dimensions. In vacuum $\phi_0 = c \hat{n}$, (\hat{n} unit vector in x, y space). Example, n points radially $\hat{n} = \hat{\rho}$ in polar coordinates (ρ, α) . Internal space ϕ_1, ϕ_2 . For $|\mathbf{r}| >$ soliton radius $\phi_s(x)$ is a map from $R^{d-1} - \{|\mathbf{r}| < \text{soliton radius}\}$ to manifold of possible vacua.

Ansatz for soliton in U(1) notation:

$$\begin{aligned}\phi(\rho, \alpha) &= f(\rho) e^{i\alpha} \\ U &= \int d^2x \frac{1}{2} (\nabla\phi)^* \cdot (\nabla\phi) + \frac{1}{4} \lambda (\phi^* \phi - c^2)^2 \\ &= \int d\rho 2\pi\rho \left(\frac{1}{2} (f')^2 + \frac{1}{2\rho^2} + \frac{1}{4} \lambda (f^2 - c^2)^2 \right)\end{aligned}$$

Unfortunately this does not work, for

$$\int d\rho \rho \frac{1}{\rho^2} = \infty$$

Diverges at infinity \Rightarrow infinite energy.

Gauged O(2) model in $d = 2 + 1$.

$$\mathcal{L} = (D_\mu\phi)^* (D^\mu\phi) - \frac{1}{4} \lambda (\phi^* \phi - c^2)^2 - \frac{1}{4g^2} F_{\mu\nu} F^{\mu\nu}$$

$$D_\mu\phi = (\partial_\mu - i A_\mu)\phi$$

Ansatz: $\phi_s(\rho, \alpha) = c e^{i\alpha} \chi(\rho)$. $\chi(0) = 0$, $\chi(\infty) = 1$.

$$A_0 = 0, \quad \mathbf{A}(\rho, \alpha) = \hat{\alpha} A_\alpha(\rho)$$

$$\mathbf{D}\phi = (\nabla - i\mathbf{A})\phi = \left(\hat{\rho} \frac{\partial}{\partial\rho} + \frac{1}{\rho} \hat{\alpha} \frac{\partial}{\partial\alpha} - i\mathbf{A} \right) \phi = c e^{i\alpha} \left[\hat{\rho} \chi' + i \hat{\alpha} \left(\frac{1}{\rho} - A_\alpha \right) \right] \chi$$

$$\nabla \times \mathbf{A} = \frac{1}{\rho} \partial_\rho (\rho A_\alpha)$$

$$\begin{aligned}U &= \int d^2x \left[(\mathbf{D}\phi)^* \cdot (\mathbf{D}\phi) + \frac{1}{4} \lambda (\phi^* \phi - c^2)^2 + \frac{1}{2g^2} (\nabla \times \mathbf{A})^2 \right] \\ &= 2\pi c^2 \int_0^\infty d\rho \rho \left[\chi'^2 + \left(\frac{1}{\rho} - A_\alpha \right)^2 \chi^2 + \frac{1}{4} \lambda c^2 (1 - \chi^2)^2 + \frac{1}{2g^2} \frac{1}{\rho^2} (\partial_\rho (\rho A_\alpha))^2 \right]\end{aligned}$$

It is easy to see that minimum exists and $0 < U < \infty$, so this solution exists. In Landau Ginsburg model it describes flux tube in superconductor. Magnetic flux through tube:

$$\int_0^{2\pi} A_{em} \rho d\alpha = \frac{1}{2e} \int A \rho d\alpha = \frac{2\pi}{2e}$$

$$g \rightarrow 2e, A_\mu \rightarrow 2e A_{em}.$$

Solitons in $d = 3 + 1$? Yes.

Gauged O(3) model in 3 + 1 dimensions with spontaneous symmetry breaking has solitons too.

First spontaneous symmetry breaking in O(3) model.

In ungauged model O(3) was spontaneously broken to O(2) and there are 2 Goldstone bosons. In gauged model this means that 2 of the 3 vector fields become massive, one remains massless, called EM field, one can make ansatz and find stable soliton as in O(2) case. This soliton has magnetic charge $q = 4\pi/e$.